

# Astronomy Equations

## Chapter 1 Measuring Distance

$parallax = \frac{360}{2\pi} \times \frac{baseline}{distance}$	$diameter = distance \times \frac{angular\ diameter}{57.3}$	$distance = baseline \times \frac{57.3}{parallax}$
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## Chapter 2 Copernican Revolution

$p^2 (in\ earth\ years) = a^3 (in\ Au's)$	$F = \frac{Gm_1m_2}{r^2}$	$F = ma$
$p^2 (in\ earth\ years) = \frac{a^3 (in\ Au's)}{M_{total}(in\ solar\ units)}$	$a = \frac{v^2}{r}$	$v_{escape} = \sqrt{2 \frac{GM}{r}}$
$V = \sqrt{\frac{GM}{r}}$		

## Chapter 3 Radiation

$frequency = \frac{1}{period}$	$Velocity = \frac{wavelength}{period}$	$Velocity = wavelength \times frequency$
$Wavelength\ of\ peak\ emission \propto \frac{1}{temp}$	$total\ energy\ emission \propto temp^4$	$\lambda_{max} = \frac{2.9mm}{T}$
$F = \sigma T^4$	$L = 4\pi\sigma R^2 T^4$ OR $luminosity \propto radius(solar\ radii)^2 \times temperature(units\ of\ 5800K)^4$	$\frac{Apparent\ wavelength}{true\ wavelength} = \frac{true\ frequency}{apparent\ frequency}$
$\frac{Apparent\ wavelength}{true\ wavelength} = 1 + \frac{recession\ velocity}{wave\ speed}$	$\frac{Apparent\ wavelength}{true\ wavelength} = \frac{recession\ velocity}{wave\ speed}$	$\frac{change\ in\ wavelength}{true\ wavelength} = \frac{recession\ velocity}{wavespeed}$
$\frac{recession\ velocity}{wavespeed, c} = \frac{change\ in\ wavelength}{true\ wavelength}$		

## Chapter 4 Spectroscopy

$E = hf$	$E_n = 13.6(1 - \frac{1}{n^2})(ev)$
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## Chapter 5 Telescopes

$Angular\ resolution(arcsec) = 0.25 \frac{wavelength(\mu m)}{diameter(m)}$
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## Chapter 6 The solar system

$Linear\ momentum = mass \times velocity$	$Angular\ momentum \propto mass \times rotation\ rate \times radius^2$
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# Astronomy Equations

## Chapter 7 Earth

$Scattering\ by\ dust \propto \frac{1}{wavelength}$	$Scattering\ by\ molecules \propto \frac{1}{wavelength^4}$	$Fraction\ of\ material\ remaining = \frac{t}{2}$
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## Chapter 8 The Moon and Mercury

$Avg.\ molecular\ speed(\frac{km}{s}) = 0.157 \sqrt{\frac{gas\ temp.(k)}{molecular\ mass}}$	$Escape\ speed(\frac{km}{s}) = 11.2 \sqrt{\frac{mass\ of\ body}{radius\ of\ mass}}$
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## Chapter 16 The Sun

$\frac{Solar\ luminosity}{solar\ mass} = 2 \times 10^{-4} \frac{W}{kg}$	$E = mc^2$
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## Chapter 17 The Sun

$Distance(parsecs) = \frac{1}{parallax(arcseconds)}$	$Apparent\ brightness(energy\ flux) \propto \frac{luminosity}{distance^2}$	$Apparent\ magnitude - absolute\ magnitude = 5 \log_{10} \frac{distance}{10pc}$
$L(solar\ units) = 10^{-(M-4.83)/2.5}$	$R = \frac{\sqrt{L}}{T^2}$	$Stellar\ lifetime \propto \frac{stellar\ mass}{stellar\ luminosity}$
$Stellar\ lifetime \propto \frac{1}{(stellar\ mass)^3}$		

## Chapter 22 Neutron Stars and Black Holes

$Deflection(arc\ sec.) = 1.75 \frac{M(solar\ masses)}{R(solar\ radii)}$
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## Chapter 23 The Milky Way Galaxy

$Total\ mass\ (solar\ masses) = \frac{orbital\ size\ (Au)^3}{orbital\ period\ (years)^3}$
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## Chapter 24 Galaxies

$Recessional\ velocity = H_0 \times distance$	$Redshift = \frac{observed\ wave\ length - true\ wavelength}{true\ wavelength} = \frac{recessional\ velocity\ (v)}{speed\ of\ light\ (c)}$
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## Chapter 25 Galaxies and Dark Matter

$Time = \frac{distance}{velocity} = \frac{1}{H_0}$	$\frac{1\ Joule}{(3 \times 10^8 \frac{m}{s})^2} = 1.1 \times 10^{-17} kg$
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## Chapter 27 The Early Universe

$C = \lambda f$
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